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# Absorption of a randomly accelerated particle: gambler's ruin in a different game

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**Abstract.** We consider a particle which is randomly accelerated by Gaussian white noise on the line 0 < x < 1, with absorbing boundaries at x = 0, 1. Denoting the initial position and velocity of the particle by  $x_0$  and  $v_0$  and solving a Fokker–Planck-type equation, we derive the exact probabilities  $q_0(x_0, v_0), q_1(x_0, v_0)$  of absorption at x = 0, 1, respectively. The results are in excellent agreement with computer simulations.

A well known topic in random walk theory [1] is the problem of the 'gambler's ruin'. Initially, the gambler has an amount of money  $x_0$  and the bank the amount  $1-x_0$ . The gambler flips a coin repeatedly, randomly winning or losing the increment  $\epsilon$ . The game ends when the gambler's funds reach 0 or 1. The problem is to compute the probability  $q_0(x_0)$  that the gambler loses everything.

The problem is easily solved. Since  $q_0(x_0) = \frac{1}{2}[q_0(x_0 + \epsilon) + q_0(x_0 - \epsilon)]$ ,

$$\frac{d^2 q_0(x_0)}{dx_0^2} = 0 \tag{1}$$

in the limit  $\epsilon \to 0$ . From equation (1) and the boundary conditions  $q_0(0) = 1$ ,  $q_0(1) = 0$ ,

$$q_0(x_0) = 1 - x_0. (2)$$

As the starting capital increases from 0 to 1, the probability of the gambler's ruin decreases from 1 to 0.

Instead of the gambling scenario one could equally well imagine a particle making a random walk with infinitesimal steps  $\pm \epsilon$  on the *x*-axis, with initial position  $0 < x_0 < 1$ . Over the course of time the particle eventually arrives at x = 0 or 1. The quantities  $q_0(x_0)$  in equation (2) and  $q_1(x_0) = 1 - q_0(x_0)$  represent the probabilities that the particle first reaches the edge of the interval at x = 0 and 1, respectively. Alternatively, we could impose absorbing boundary conditions and interpret  $q_0(x_0)$  and  $q_1(x_0)$  as the probabilities of absorption at x = 0 and at x = 1.

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In this paper we also consider a particle on the finite interval 0 < x < 1, but we assume that the changes in the *velocity* rather than the position of the particle are random. The particle moves according to the Langevin equation

$$\frac{\mathrm{d}^2 x}{\mathrm{d}t^2} = \eta(t) \tag{3}$$

where the acceleration  $\eta(t)$  has the form of Gaussian white noise, with

$$\langle \eta(t) \rangle = 0 \qquad \langle \eta(t_1)\eta(t_2) \rangle = 2\delta(t_1 - t_2). \tag{4}$$

Imposing absorbing boundary conditions, we derive the probabilities  $q_0(x_0, v_0)$ ,  $q_1(x_0, v_0)$  of absorption at x = 0 and at x = 1, respectively, as functions of the initial position and velocity.

The quantity  $q_0(x_0, v_0)$  can also be interpreted as the probability of a gambler's ruin, but the game is different. The gambler has an amount of money x(t) at time t and the bank the amount 1 - x(t). Money is transferred from the bank to the gambler at a rate v = dx/dt, which may be positive or negative. At regular infinitesimal intervals the gambler flips a coin, randomly increasing or decreasing the rate v by the increment  $\Delta$ . The game ends when x reaches 0 or 1. The quantity  $q_0(x_0, v_0)$  is the probability that a gambler with initial conditions  $x_0, v_0$  loses everything.

In the case of a random walk on the x-axis, the probability density  $P(x, x_0, t)$  at time t of a particle which is initially at  $x_0$  obeys the diffusion equation. For a particle which is randomly accelerated according to equations (3) and (4), the probability density  $P(x, v; x_0, v_0; t)$  in the phase space (x, v) satisfies the Fokker–Planck equation [2]

$$\left(\frac{\partial}{\partial t} + v\frac{\partial}{\partial x} - \frac{\partial^2}{\partial v^2}\right)P(x, v; x_0, v_0; t) = 0$$
(5)

corresponding to diffusion of the velocity, with the initial condition

$$P(x, v; x_0, v_0; 0) = \delta(x - x_0)\delta(v - v_0).$$
(6)

In analogy with the discussion leading to differential equation (1) for  $q_0(x_0)$ , let us consider a discrete dynamics in which the velocity v changes by  $\pm \Delta$  with equal probability at time intervals  $\tau$ . For this dynamics

$$P(x, v; x_0, v_0; t) = \frac{1}{2} [P(x - v\tau, v + \Delta; x_0, v_0; t - \tau) + P(x - v\tau, v - \Delta; x_0, v_0; t - \tau)]$$
(7)

$$q_0(x_0, v_0) = \frac{1}{2} [q_0(x_0 + v_0\tau, v_0 + \Delta) + q_0(x_0 + v_0\tau, v_0 - \Delta)].$$
(8)

Expanding equations (7) and (8) in  $\tau$  and  $\Delta$ , dividing by  $\tau$ , and taking the limit  $\tau = \frac{1}{2}\Delta^2 \rightarrow 0$  gives us a 'poor man's' derivation of the Fokker–Planck equation (5) and the corresponding differential equation

$$\left(v_0\frac{\partial}{\partial x_0} + \frac{\partial^2}{\partial v_0^2}\right)q_0(x_0, v_0) = 0$$
(9)

for the probability of absorption at x = 0.

To solve equation (9) with the absorbing boundary condition

 $q_0(0, v_0) = 1 \qquad v_0 < 0 \tag{10}$ 

and the requirements

$$q_0(x_0, v_0) = q_1(1 - x_0, -v_0) \tag{11}$$

$$q_0(x_0, v_0) + q_1(x_0, v_0) = 1$$
(12)

of reflection symmetry and total probability equal to 1, we first make the substitution

$$\psi(x,v) = q_0(x,-v) - \frac{1}{2}.$$
(13)

Expressed in terms of  $\psi(x, v)$ , equations (9)–(12) take the form

$$\left(v\frac{\partial}{\partial x} - \frac{\partial^2}{\partial v^2}\right)\psi(x,v) = 0 \tag{14}$$

$$\psi(0,v) = \frac{1}{2} \qquad v > 0 \tag{15}$$

$$\psi(x, v) = -\psi(1 - x, -v).$$
(16)

Masoliver and Porrà [3] have shown how certain Fokker–Planck-type equations on the finite interval 0 < x < 1 can be solved exactly. They derived an exact result for the average time  $T(x_0, v_0)$  a randomly accelerated particle with initial conditions  $x_0, v_0$  takes to reach a boundary of the interval. The probability that the particle has not yet reached a boundary after a time *t* decays as  $e^{-Et}$ , as discussed by Burkhardt [4]. He obtained *E* numerically with an approach similar to [3] and related it to the confinement free energy of a semiflexible polymer in a tube. In another application inspired by [3], Burkhardt *et al* [5] calculated the equilibrium distribution function P(x, v) of a randomly accelerated particle on the line 0 < x < 1 undergoing inelastic collisions at the boundaries [6].

The function  $\psi(x, v)$  satisfies the same steady-state Fokker–Planck equation (14) as the quantity P(x, v) considered in [5] and has the same Green's function solution

$$\psi(x,v) = \frac{v^{1/2}}{3x} \int_0^\infty du \, u^{3/2} e^{-(v^3 + u^3)/9x} I_{-1/3}\left(\frac{2v^{3/2}u^{3/2}}{9x}\right) \psi(0,u) -\frac{1}{3^{1/3}\Gamma(\frac{2}{3})} \int_0^x dy \, \frac{e^{-v^3/9(x-y)}}{(x-y)^{2/3}} \frac{\partial\psi(y,0)}{\partial v} \quad v > 0$$
(17)

derived in [5]. Equation (17) only holds for positive v. For negative v,  $\psi(x, v)$  can be obtained from equation (17) using the antisymmetry (16) under reflection.

Equation (17) determines  $\psi(x, v)$  for all x > 0 and v > 0 from  $\psi(0, v)$  and  $\partial \psi(x, 0)/\partial v$ . The first of these functions is given in equation (15). To determine the second, we set v = 0 in equation (17), which yields

$$\psi(x,0) = \frac{1}{3^{1/3}\Gamma(\frac{2}{3})} \left[ x^{-2/3} \int_0^\infty du \, u e^{-u^3/9x} \psi(0,u) - \int_0^x \frac{dy}{(x-y)^{2/3}} \, \frac{\partial\psi(y,0)}{\partial v} \right].$$
(18)

Then, substituting equation (18) in the relation  $\psi(x, 0) + \psi(1 - x, 0) = 0$ , which follows from (16), and using  $\partial \psi(y, 0) / \partial v = \partial \psi(1 - y, 0) / \partial v$ , also a consequence of (16), we obtain

$$\int_0^1 \frac{\mathrm{d}y}{|x-y|^{2/3}} \frac{\partial \psi(y,0)}{\partial v} = \int_0^\infty \mathrm{d}u \, u \left[ \frac{\mathrm{e}^{-u^3/9x}}{x^{2/3}} + \frac{\mathrm{e}^{-u^3/9(1-x)}}{(1-x)^{2/3}} \right] \psi(0,u).$$
(19)

The solution to integral equation (19), derived, following [7], in appendix B of [5], is given by

$$\frac{\partial \psi(x,0)}{\partial v} = \int_0^\infty \mathrm{d}u \, u \left[ R(x,u) + R(1-x,u) \right] \psi(0,u) \tag{20}$$

where

$$R(x,u) = \frac{1}{3^{5/6}\Gamma(\frac{1}{3})\Gamma(\frac{5}{6})} \frac{u^{1/2} e^{-u^3/9x}}{x^{7/6}(1-x)^{1/6}} {}_1F_1\left(-\frac{1}{6},\frac{5}{6},\frac{u^3(1-x)}{9x}\right)$$
(21)

and  $_1F_1(a; b; z)$  is the confluent hypergeometric function [8, 9].

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Equations (17), (20), and (21) determine  $\psi(x, v)$  for all x and v from  $\psi(0, v)$  for v > 0, which is known from the absorbing boundary condition (15). Substituting equations (15) and (21) in (20) leads to

$$\frac{\partial \psi(x,0)}{\partial v} = \frac{1}{3^{1/6} \Gamma(\frac{1}{3})} [x(1-x)]^{-1/6}$$
(22)

and from (15), (17) and (22)

$$\psi(x,v) = \frac{1}{2} - \frac{1}{2\pi} \int_0^x dy \, \frac{e^{-v^3/9(y-x)}}{(y-x)^{2/3}} \, [y(1-y)]^{-1/6}.$$
(23)

Rewriting equation (23) in terms of  $q_0(x_0, v_0)$  using (11)–(13), we obtain our main result

$$q_0(x_0, v_0) = 1 - q_0(1 - x_0, -v_0) = \frac{1}{2\pi} \int_{x_0}^1 dy \, \frac{e^{-v_0^3/9(y - x_0)}}{(y - x_0)^{2/3}} \, [y(1 - y)]^{-1/6} \qquad v_0 > 0$$
(24)

analogous to the solution (2) of the traditional gambler's ruin problem.

For  $x_0 = 1$  equation (24) reproduces the expected result  $q_0(1, v_0) = 1 - q_0(0, -v_0) = 0$ ,  $v_0 > 0$ , corresponding to the immediate absorption of a particle that is initially at either boundary with velocity directed outward from the interval 0 < x < 1. For  $x_0 = 0$  and  $v_0 = 0$  the integral in equation (24) can be evaluated, yielding

$$q_0(0, v_0) = 1 - q_0(1, -v_0) = 1 - \frac{2 \times 3^{2/3}}{\Gamma(\frac{1}{6})} v_0^{1/2} {}_1F_1(\frac{1}{6}; \frac{7}{6}; -\frac{1}{9}v_0^3) \qquad v_0 > 0$$
(25)

$$q_0(x_0,0) = 1 - q_0(1-x,0) = 1 - \frac{6\Gamma(\frac{1}{3})}{\Gamma(\frac{1}{6})^2} x_0^{1/6} {}_2F_1(\frac{1}{6},\frac{5}{6};\frac{7}{6};x_0).$$
(26)

Here  $_1F_1(a; b; z)$  and  $_2F_1(a, b; c; z)$  are the confluent and ordinary hypergeometric functions [8,9].

The probability  $q_0(x_0, v_0)$  of absorption at the origin, obtained from equation (24) by numerical integration, is shown in figure 1. The probability decreases monotonically as  $x_0$ increases with fixed  $v_0$  and as  $v_0$  increases at fixed  $x_0$ , as expected. The quantity  $q_0(x_0, v_0)$  is a non-singular function of  $(x_0, v_0)$  except at the two boundary points (0, 0) and (1, 0). The curves for  $x_0 = 0.0, 0.1, 0.3, 0.5$  become smoother near  $v_0 = 0$  as  $x_0$  increases, and for  $x_0 = 0.5, q_0(x_0, v_0) - \frac{1}{2}$  is an odd function of  $v_0$ , as implied by equations (11) and (12).

The points in figure 1 show the results of computer simulations, which clearly are in excellent agreement with the analytical results. Our simulation routine is based on the exact solution [10]

$$P_{\text{free}}(x,v;x_0,v_0;t) = \frac{\sqrt{3}}{2\pi t^2} \exp\left\{-\frac{3}{t^3} \left[(x-x_0-v_0t)(x-x_0-vt) + \frac{1}{3}(v-v_0)^2t^2\right]\right\}$$
(27)

of the Fokker–Planck equation (5) with initial condition (6) in the absence of boundaries. Trajectories with the probability distribution  $P_{\text{free}}(x_{n+1}, v_{n+1}; x_n, v_n; \Delta_{n+1})$  given by (27) are generated using the algorithm

$$x_{n+1} = x_n + v_n \Delta_{n+1} + \left(\frac{\Delta_{n+1}^3}{6}\right)^{1/2} \left(s_{n+1} + \sqrt{3} r_{n+1}\right)$$
(28)

$$v_{n+1} = v_n + (2\Delta_{n+1})^{1/2} r_{n+1}$$
<sup>(29)</sup>



**Figure 1.** The probability  $q_0(x_0, v_0)$  of absorption at the origin. The full curves show the exact result given in equation (24). The points are the results of our computer simulations. The data points have a statistical uncertainty  $\pm \delta q_0$  with  $|\delta q_0| \lesssim 0.001$ .

where  $x_n$  and  $v_n$  are the position and velocity of the particle at time  $t_n$ , and  $\Delta_{n+1} = t_{n+1} - t_n$ . The quantities  $r_n$  and  $s_n$  are independent Gaussian random numbers such that

$$\langle r_n \rangle = \langle s_n \rangle = 0 \qquad \langle r_n^2 \rangle = \langle s_n^2 \rangle = 1.$$
 (30)

In the absence of boundaries there is no time-step error in the algorithm, i.e. the  $\Delta_n$  may be chosen arbitrarily. Close to boundaries small time steps are needed.

To derive a quantitative criterion for an acceptable time step, we begin with the averages

$$\langle x(t) \rangle = x_0 + v_0 t \qquad \langle [x(t) - \langle x(t) \rangle]^2 \rangle = \frac{2}{3} t^3$$
(31)

implied by the distribution function (27). At time *t* the particle coordinate *x* has a Gaussian distribution, with a maximum at  $x = x_0 + v_0 t$  and the root-mean-square width  $\left(\frac{2}{3}t^3\right)^{1/2}$ . The effect of the boundaries on the propagation is negligible if the Gaussian peak lies almost entirely within the interval 0 < x < 1. This is certainly the case if, say,

$$0 < x_0 + v_0 t \pm 5t^{3/2} < 1.$$
(32)

Over the range of velocities encountered in our simulations, any t which satisfies the simpler, more stringent condition

$$t < \frac{1}{10}x_0(1 - x_0) \tag{33}$$

also satisfies (32).

Keeping inequality (33) in mind, we performed our simulations with the time step

$$\Delta_{n+1} = 10^{-5} + 10^{-1} x_n (1 - x_n). \tag{34}$$

The time step decreases as the particle approaches the boundary and has the minimum value  $10^{-5}$ . It is necessary to have a small non-zero minimum value. Otherwise the particle never

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arrives at the boundaries. Our results for the absorption probability  $q_0(x_0, v_0)$  are averages based on  $10^5$  trajectories for each set of initial conditions  $x_0, v_0$ .

Finally, we note that  $q_0(x_0, v_0)$  may be derived from another general Green's function solution<sup>†</sup> of the Fokker–Planck equation (14),

$$\psi(x,v) = \frac{v^{1/2}}{3x} \int_0^\infty du \, u^{3/2} e^{-(v^3 + u^3)/9x} \, I_{1/3}\left(\frac{2v^{3/2}u^{3/2}}{9x}\right) \psi(0,u) + \frac{v}{3^{2/3}\Gamma(\frac{1}{3})} \int_0^x dy \, \frac{e^{-v^3/9(x-y)}}{(x-y)^{4/3}} \, \psi(y,0) \qquad v > 0$$
(35)

different from (17). By substituting equation (35) in (14), one can check that the Fokker–Planck equation is indeed satisfied. On the lines x = 0 and v = 0 equation (35) reduces to the identities  $\psi(0, v) = \psi(0, v)$  and  $\psi(x, 0) = \psi(x, 0)$ , respectively.

Equation (35) determines  $\psi(x, v)$  for all x > 0 and v > 0 from  $\psi(0, v), v > 0$  and  $\psi(x, 0)$ . The first of these functions is given in equation (15). To determine the second one, we differentiate equation (35) with respect to v and then set v = 0, which yields

$$\frac{\partial \psi(x,0)}{\partial v} = \frac{1}{3^{2/3} \Gamma(\frac{1}{3})} \bigg[ x^{-4/3} \int_0^\infty du \, u^2 e^{-u^3/9x} \psi(0,u) - 3x^{-1/3} \psi(0,0) \\ -3 \int_0^x \frac{dy}{(x-y)^{1/3}} \, \frac{\partial \psi(y,0)}{\partial y} \bigg].$$
(36)

For the absorbing boundary condition (15) the first two terms on the right-hand side of (36) cancel. Substituting equation (36) in the relation  $\partial \psi(x, 0)/\partial v - \partial \psi(1-x, 0)/\partial v = 0$ , which follows from (16), using the invariance of  $\partial \psi(y, 0)/\partial y$  under  $y \to 1-y$ , and integrating with respect to x yields

$$\int_0^1 dy |x - y|^{2/3} \frac{\partial \psi(y, 0)}{\partial y} = \text{constant.}$$
(37)

The function  $\psi(x, 0)$  given in equations (13) and (26) satisfies equation (37). Substituting this  $\psi(x, 0)$  and  $\psi(0, v) = \frac{1}{2}$  into equation (35), integrating, and using (11)–(13), we obtain

$$q_0(x_0, v_0) = 1 - q_0(1 - x_0, -v_0)$$
  
=  $\frac{2 \times 3^{1/3} v}{\Gamma(\frac{1}{6})^2} \int_{x_0}^1 dy \, \frac{e^{-v_0^3/9(y - x_0)}}{(y - x_0)^{4/3}} \, (1 - y)^{1/6} \, {}_2F_1(\frac{1}{6}, \frac{5}{6}; \frac{7}{6}; 1 - y) \qquad v_0 > 0.$  (38)

With the help of the identity

$$\int_0^x \mathrm{d}y \, \frac{\mathrm{e}^{-\nu^3/9(x-y)}}{(x-y)^{2/3}} \, f(y) = \frac{\nu}{3^{2/3} \Gamma(\frac{1}{3})} \int_0^x \mathrm{d}z \, \frac{\mathrm{e}^{-\nu^3/9(x-z)}}{(x-z)^{4/3}} \int_0^z \mathrm{d}y \, \frac{f(y)}{(z-y)^{2/3}} \tag{39}$$

for arbitrary f(y), one can convert expression (38) for  $q_0(x_0, v_0)$  into the simpler form (24).

The second of the two Green function solutions (17), (35) looks simpler than the first, since no derivatives of  $\psi$  appear on the right-hand side, but our main result (24) for  $q_0(x_0, v_0)$  is obtained more easily from (17).

<sup>&</sup>lt;sup>†</sup> This solution may be derived by slightly modifying the derivation in appendix A of [5]. Setting v = 0 in equation (A3) of [5], solving for W(s), and reinserting the result in (A3) with  $v \neq 0$  yields the Laplace transform of the new solution (35).

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